Implication of 126 GeV Higgs for Planck scale physics

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Higgs was discovered at $M_{\rm H}$ =126 GeV

No evidence of "new" physics @ ATLAS,CMS & LHCb

What is the implication of these two?

together with some phenomena beyond SM (v oscillation, Baryon asymmetry, Dark matter)

3 major hints towards the physics beyond SM





which is consistent with SM (3.2±0.2)×10⁻⁹.

(3) But we also know that SM is not sufficient to explain neutrino oscillation Baryon asymmetry dark matter

(also a big hint from Cosmology)

Most investigations of physics beyond the SM have been based on

"the central dogma" of particle physics GUT \rightarrow hierarchy problem \rightarrow TeV SUSY etc.

i.e. Unification below the Planck scale requires large symmetry enhancement at TeV scale.

It may be a good time to reconsider the basic strategy (central dogma) toward the physics BSM.

Hierarchy problem

Naturalness (Hierarchy problem)

$$\delta V(\phi) = \frac{1}{2} \int \frac{d^4k}{(2\pi)^4} \operatorname{Str} \log(k^2 + M^2(\phi)) \qquad \text{Quadratic divergence}$$

$$= \frac{\Lambda^2}{32\pi^2} \operatorname{STr} M^2(\phi) + \operatorname{STr} \frac{M^4(\phi)}{64\pi^2} (\ln(M^2/\Lambda^2) - 1/2)$$

- $\mathrm{STr}M^2(\phi) \neq 0$
- Quadratic divergence in Higgs mass term

 $\mathrm{STr}M^2(\phi) = 0$

Cancellation of Quadratic divergence (supersymmetry etc.)

Bardeen (1995 @ Ontake summer school)

Standard model is classically scale invariant if Higgs mass term is absent.

$$T^{\mu}_{\mu} = 0$$

Quantum anomaly breaks the invariance (if not conformal)

$$T^{\mu}_{\mu} = \beta(\lambda_i)\mathcal{O}_i$$

But then, we cannot forbid the quadratically divergent mass term $T^{\mu}_{\mu} = \beta(\lambda_i)\mathcal{O}_i + const. \ \Lambda^2 \bar{h}h$

Bardeen argued that it should be

$$T^{\mu}_{\mu} = \beta(\lambda_i)O_i + \delta m^2 \bar{h}h$$

$$\delta m^2 = const. \times m^2 \neq const. \times \Lambda^2$$

Is quadratic divergence the issue of hierarchy problem?

NO

Bardeen(1995)

H Aoki, SI PRD(2012)

There are 3 different types of divergences

- 1. Quadratic divergences Λ^2
- 2. Logarithmic divergences $m^2 \log \Lambda$

3. Logarithmic but looks like quadratic $M^2 \log \Lambda$

(1) Quadratic divergence can be simply subtracted, so it gives a boundary condition at UV cut off Λ .

 \rightarrow If massless at Λ , it continues to be so in the IR theory.

(2) Logarithmic divergence gives a multiplicative renormalization.
 No Higgs mass term is generated if it is absent at UV scale.

$$\frac{dm^2}{dt} = \frac{m^2}{16\pi^2} \left(12\lambda + 6Y_t^2 - \frac{9}{2}g^2 - \frac{3}{2}g_1^2 \right) + \frac{M^2}{8\pi^2} \lambda_{mix}$$
(3) If SM is coupled with a massive particle with mass M, logarithmic divergences give a correction to m as

$$\delta m^2 = \frac{\lambda_{mix}M^2}{16\pi^2} \log(\Lambda^2/m^2)$$

In order to solve the "hierarchy problem" without a special cancellation like supersymmetry, we need to control

 (a) "quadratic divergence" → correct boundary condition at Planck The most natural b.c. is NO MASS TERMS at Planck (= classical conformal invariance)

(b) "large logarithmic divergence" by mixing with a large mass M No intermediate scales between EW (or TeV) and Planck

"Classical conformal theory with no intermediate scale" can be an alternative solution to hierarchy problem.

Bardeen (95) Shaposhnikov (07) Meissner Nicolai (07) SI, Okada,Orikasa (09) 9

Stability of Vacuum

Another Hint of 126 GeV Higgs mass is Stability bound of the Higgs quartic coupling



 $m_H = 126 \text{ GeV}$

v = 246 GeV

 $V_{\rm eff}(h) = \frac{\lambda_{\rm eff}(h)}{\Lambda} h^4$ RGE improved effective potential for large field (h >> v)

$$\text{RGE @1-loop} \quad \frac{d\lambda_H}{dt} = \frac{1}{16\pi} \left(24\lambda^2 \left[-6Y_t^4 + \frac{9}{8}g^4 + \frac{3}{8}g_1^4 + \cdots \right] \right)$$
 Already known

It is related to Higgs mass as $M_h^2 = 2\lambda v^2$

Higgs mass controls the behavior of Higgs potential at large values of h.

This gives two bounds for Higgs mass 800 (1) The quartic coupling does not blow up until UV cut-off. 600 $M_{\rm H}~({\rm GeV})$ M < 180 GeV (triviality bound) (2) The quartic coupling does not become negative 400 until UV cut-off. (Stability bound) 200

M = 125 GeV Higgs is very close to the stability bound.



Why stability bound is important for Planck scale physics?



Alkhin, Djouadi, Moch (12)



If this



 $\lambda(\Lambda_0) = \beta_\lambda(\Lambda_0) = 0$

Direct window to Planck scale

M.Shaposhnikov (07)

Emergence of Higgs potential

at the Planck scale

Indication on the Higgs potential

$$V = -\mu^2 |H|^2 + \lambda (|H|^2)^2$$

Hierarchy (classical conformality) Stability vanish at Planck

LHC data implies that Higgs has a flat potential V(H)=0 at Planck.

How can we achieve EW symmetry breaking from V(H)=0 potential at Planck?

Everything should be radiatively generated.

Two mechanisms of symmetry breaking

(1) SSB by negative mass term

$$V = \frac{\lambda}{4}h^4 + \frac{\mu^2}{2}h^2 \quad (\mu^2 < 0) \longrightarrow \quad m_h^2 = 2|\mu^2| = 2\lambda \langle h \rangle^2$$

(2) Coleman-Weinberg mechanism (radiative breaking)

$$V_{eff} = \frac{\lambda h^4}{4} + Bh^4 \left(\ln \left(\frac{h^2}{\langle h \rangle^2} \right) - \frac{25}{6} \right) \qquad B = \frac{3}{64\pi^2} \left(3\lambda^2 + \frac{3g^4 + 2g^2g'^2 + g'^4}{16} - Y_t^4 \right)$$

tree 1-loop



Higgs mass is given by

$$V''|_{\langle h\rangle} = m_h^2 = 8B\langle h\rangle^2 = \frac{6}{11}\lambda\langle h\rangle^2$$

図 2.1 improve されていない有効ポテンシャル

(2') RG improved CW mechanisms



cf. Dimensional transmutation in QCD

$$\Lambda_{QCD} = M_{UV} \exp\left(-\frac{2\pi}{b_0 \alpha_s(M_{UV})}\right)$$



Model: (B-L) extension of SM with Right Handed Neutrinos

	$\mathrm{SU}(3)_c$	$\mathrm{SU}(2)_L$	$\mathrm{U}(1)_Y$	$U(1)_{B-L}$
q_L^i	3	2	+1/6	+1/3
u_R^i	3	1	+2/3	+1/3
d_R^i	3	1	-1/3	+1/3
ℓ_L^i	1	2	-1/2	-1
$\left(\nu_{R}^{i} \right)$	1	1	0	-1
$e_R^{\overline{i}}$	1	1	-1	-1
H	1	2	-1/2	0
(Φ)	1	1	0	+2

N Okada, Y Orikasa, SI PLB676(09)81, PRD80(09)115007 PRD83(11)093011

 ${f H}$ Higgs doublet ${f \Phi}$ B-L sector scalar field

Soo cow mochanicm

B-L is the only anomaly free global symmetry in SM.
[U(1)_{B-L}]³ is anomaly free if we have right handed fermion.
B-L gauge symmetry is broken by vev of an additional scalar.

$$\mathcal{L} \supset -Y_D^{ij} \overline{\nu_R^i} H^{\dagger} \ell_L^j - \frac{1}{2} Y_N^i \Phi \overline{\nu_R^{ic}} \nu_R^i + \text{h.c.}, \qquad \begin{pmatrix} 0 & m \\ m & M_N \end{pmatrix} \longrightarrow m_{\nu} = \frac{m^2}{M_N}$$
$$m = Y_D \langle H \rangle \qquad M_N = Y_N \langle \phi \rangle$$

B-L can be broken by CW mechanism at TeV.

$$V(\Phi, H)|_{UV} = \lambda_{\Phi} (\Phi^{\dagger} \Phi)^{2}$$

$$M_{B-L} \sim M_{Planck} \exp(-\frac{\lambda_{\phi}}{b})$$

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How about EWSB ?



Can the small scalar mixing be realized naturally?

Radiatively generated scalar mixing λ_{mix} in V(H) $V(H) = \lambda_H H^4 + \lambda_{mix} \Phi^2 H^2 \rightarrow EWSB$

YES

Why is small negative λ_{mix} generated?

$$\frac{d\lambda_{mix}}{dt} = \frac{1}{16\pi^2} \left[\lambda_{mix} \left(\begin{array}{c} & \ast \ast \ast \ast \ast \ast \end{array} \right) + 12g_{mix}^2 g_{B-L}^2 \right] \right]$$

$$\frac{dg_{mix}}{dt} \sim \frac{2}{3\pi^2} g_{B-L} g_Y^2 + g_{mix} \left(\ast \ast \ast \ast \ast \ast \right)$$

$$\bigcup \left(1 \right)_{-} \left(\begin{array}{c} g_{mix} & \bigcup (1)_{-} g_{mix} \\ & \swarrow & \swarrow & \bigcup (1)_{-} g_{mix} \\ & & \swarrow & \swarrow & -g_{mix}^2 g_{B-L}^2 \\ & & \lambda_{mix} \propto -g_{mix}^2 g_{B-L}^2 \propto -g_{B-L}^4 g_Y^4 \right]$$

$$V(H) = \lambda_H H^4 + \lambda_{mix} \Phi^2 H^2$$

$$H \rangle = \sqrt{\frac{-\lambda_{mix}}{\lambda_H}} M_{B-L}$$

$$\sim c \frac{\alpha_{B-L} \alpha_Y}{\sqrt{\lambda_H}} M_{B-L}$$

$$\int_{\mathbb{T}}^{-0.001} \int_{\mathbb{T}}^{0.000} \int_{\mathbb{T}}^{10.001} Running of Scalar mixing$$

$$\lambda_{mix} \propto -g_{B-L}^4 g_Y^4$$

$$\int_{\mathbb{T}}^{0.000} \int_{10^5 \text{ ind}}^{10^8 \text{ ind}} I0^{11} \text{ ind}^{14}$$

RGE scale [GeV]

10¹⁷

Prediction of the model

In order to realize EWSB at 246 GeV, B-L scale must be around TeV (for a typical value of α_{B-L}).



Stability bound in TeV scale B-L model

$$\frac{d\lambda_{H}}{dt} = \frac{1}{16\pi} \left(24\lambda^{2} - 6Y_{t}^{4} + \frac{9}{8}g^{4} + \frac{3}{8}g_{1}^{4} + \frac{3}{4}g^{2}g_{1}^{2} + \frac{3}{4}(g^{2} + g_{1}^{2})g_{mix}^{2} + \cdots \right)$$
An extra positive term is added
$$\int_{128.0}^{129.0} \int_{128.0}^{128.0} \int_{127.5}^{128.0} \int_{127.5}^{127.0} \int_{126.5}^{127.0} \int_{0.000}^{127.5} \int$$

Summary

•126 GeV Higgs = border of the stability bound of SM vacuum.

- → Direct window to Planck scale → Flat Higgs potential @Planck Hint for the origin of Higgs in string theory
- Occam's razor scenario beyond SM

"Classically conformal B-L model" is proposed

- (1) it can solve hierarchy problem
- (2) it can explain why B-L breaking scale is around TeV.
- (3) Stability bound can be lowered about 1 GeV

 $M_{\rm H} \simeq 128 \; {\rm GeV}$

(4) phenomenologically viable

Neutrino oscillation, resonant leptogenesis

(5) Highly predictive (or excludable)

Prediction

Z' around several TeV, $M_{\phi} < M_{Z'}$, Leptogenesis at TeV

Future projects

- Origin of flat Higgs potential at Planck
 Hierarchy problem & M_H =126 GeV
 → PNGB ? Moduli ? Gauge/Higgs ?
 - Non-susy vacua of superstring with flat V(H)
- Resonant leptogenesis Garny, Kartavsev, Hohenegger (11) Kadanoff-Baym equation (quantum Boltzman)
- Non-susy GUT at Planck scale SO(10) or E6 type Gravity or string threshold correction to RGE

Hierarchy problem in Wilsonian RG

H Aoki, SI PRD(2012)]

Critical line





Fine - tuning of the **distance from the critical line** = **Low energy mass scale**

The difference is the **choice of the coordinates** of the parameter space.

Higgs potential

$$V = -\mu^2 |H|^2 + \lambda (|H|^2)^2$$
 $|H| = \sqrt{\frac{\lambda}{2\mu^2}} \equiv \frac{v}{\sqrt{2}}$



Spontaneous Symmetry Breaking

Potential minimum

Higgs particle