

Higgs boson mass in low energy SUSY models with vector-like matters

Norimi Yokozaki (Kavli IPMU)

Endo, Hamaguchi, Iwamoto, N.Y.; 2011, 2012

Nakayama N.Y.; 2011

Sato, Tobioka, N.Y.; 2011

Endo, Hamaguchi, Ishikawa, Iwamoto, N.Y.; 2012

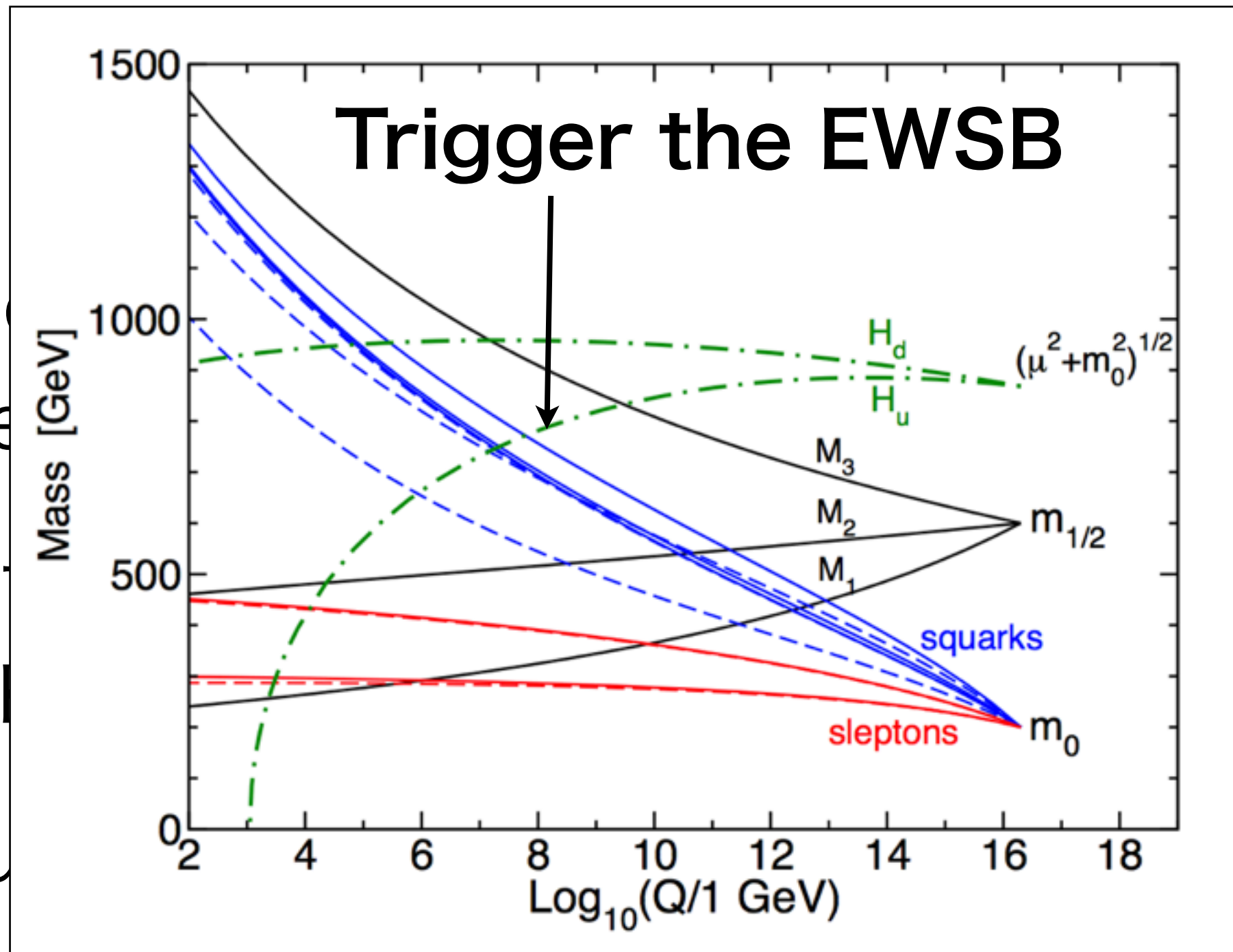
Good things of low-energy SUSY

- Radiative Electroweak symmetry breaking
- Better fine-tuning (focus point-**like** behavior can relax it more)
- Muon $g-2$ deviation ($>3\sigma$) can be explained (Sho's talk)

Good things of low-

- Rac
- bre
- Be
- bel
- Mu

explained (Sno s talk)

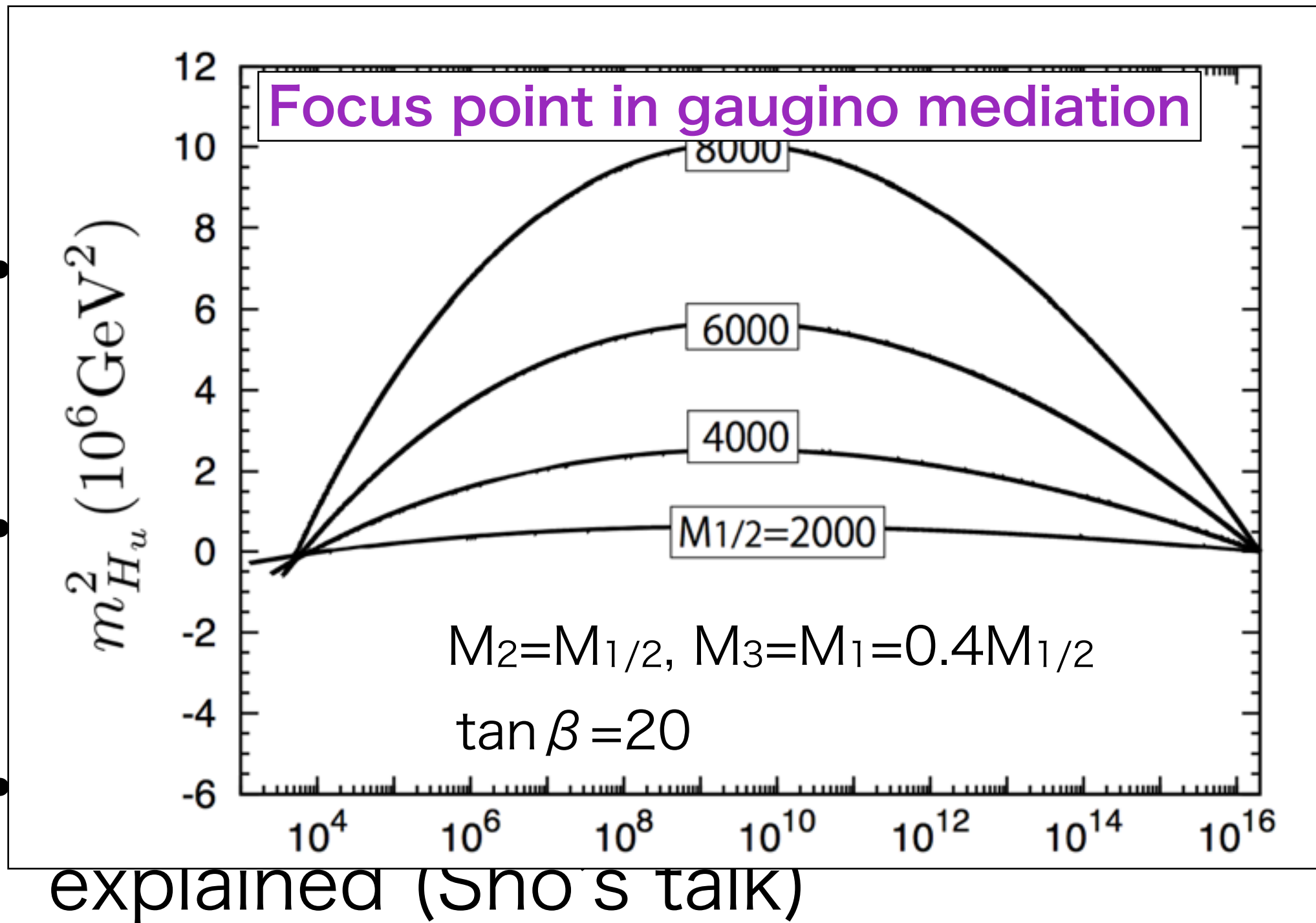


[Figures from SUSY primer]

Good things of low-energy SUSY

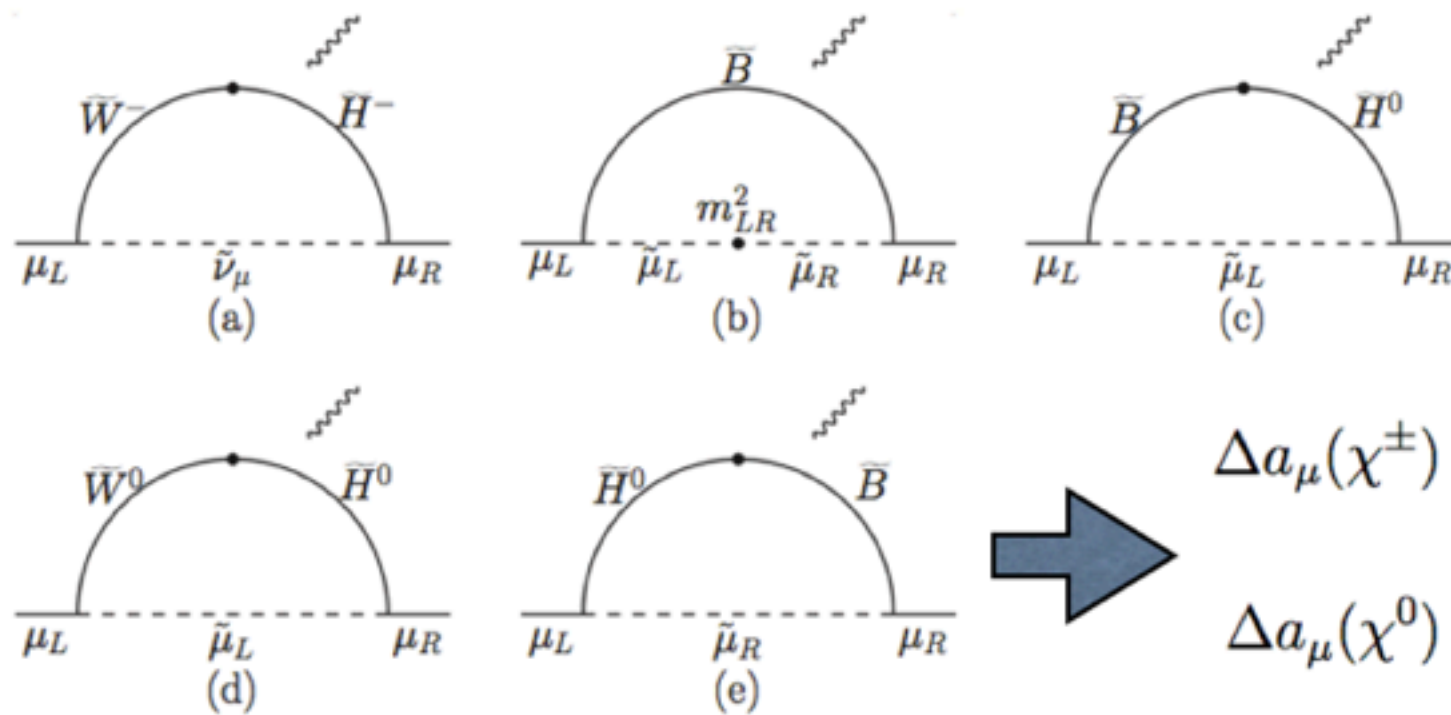
- Radiative Electroweak symmetry breaking
- Better fine-tuning (focus point-**like** behavior can relax it more)
- Muon $g-2$ deviation ($>3\sigma$) can be explained (Sho's talk)

Good things of low-



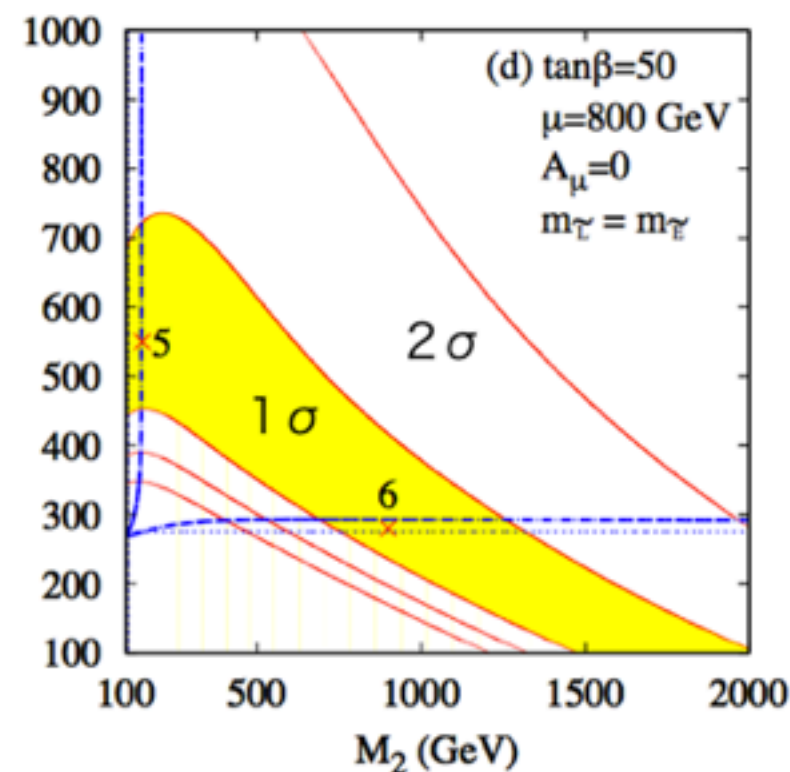
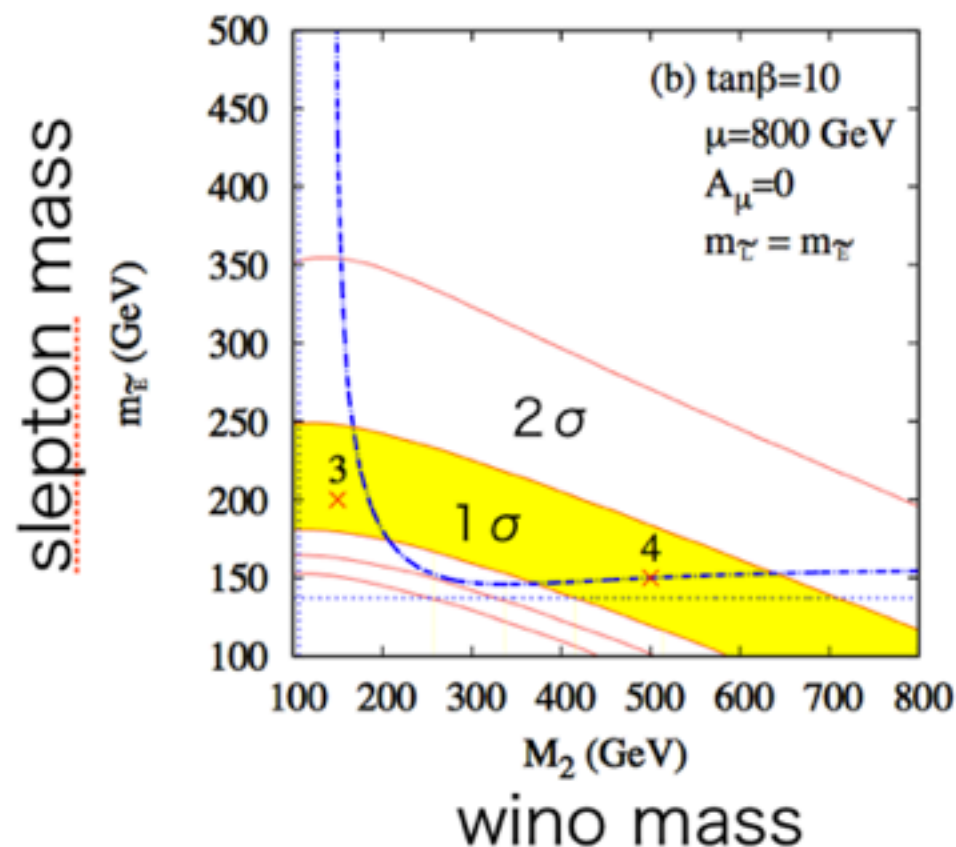
Good things of low-energy SUSY

- Radiative Electroweak symmetry breaking
- Better fine-tuning (focus point-**like** behavior can relax it more)
- Muon $g-2$ deviation ($>3\sigma$) can be explained (Sho's talk)



$$\Delta a_\mu(\chi^\pm) \simeq \frac{\alpha_2 m_\mu^2}{8\pi m_{\text{soft}}^2} \text{sgn}(\mu M_2) \tan \beta,$$

$$\Delta a_\mu(\chi^0) \simeq \frac{\alpha_Y m_\mu^2}{24\pi m_{\text{soft}}^2} \text{sgn}(\mu M_1) \tan \beta + \dots,$$



(Cho, Hagiwara, Matsumoto, Nomura, 2011)

sleptons and neutralino/chargino should be light enough

$h \gamma \gamma$ is possibly enhanced

If the SUSY particles are light, **gauge mediation models** are attractive

Gravity mediation with $O(1)$ TeV gravitino

Flavor changing neutral current
The cosmological Polonyi problem
(spoiling BBN, dilution BAU)

If the SUSY particles are light, **gauge mediation models** are attractive

Gravity mediation with $O(1)$ TeV gravitino

Flavor changing neutral current
The cosmological Polonyi problem
(spoiling BBN, dilution BAU)

However, the Higgs mass of ~ 125 GeV is
challenging

If the SUSY particles are light, **gauge mediation models** are attractive

Gravity mediation with $O(1)$ TeV gravitino

Flavor changing neutral current
The cosmological Polonyi problem
(spoiling BBN, dilution BAU)

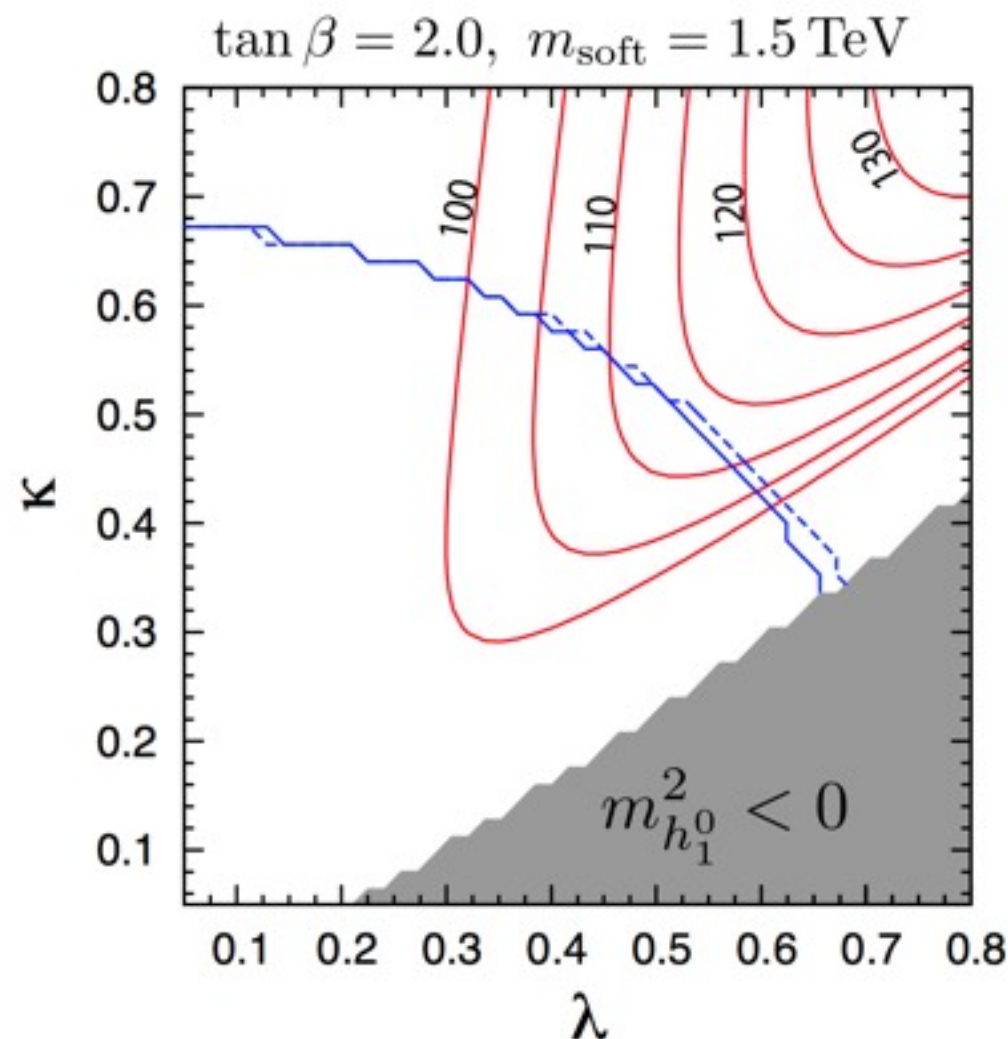
However, the Higgs mass of ~ 125 GeV is
challenging

Gauge Mediation + Extra Contribution to Higgs
mass is attractive possibility in Low-energy SUSY

Most famous example is NMSSM

NMSSM \rightarrow not good in GMSB

(Difficult to realize EWSB, Landau pole, Muon $g-2$ deviation can not be explained)



$$\lambda S H_u H_d + \kappa S^3 / 3$$

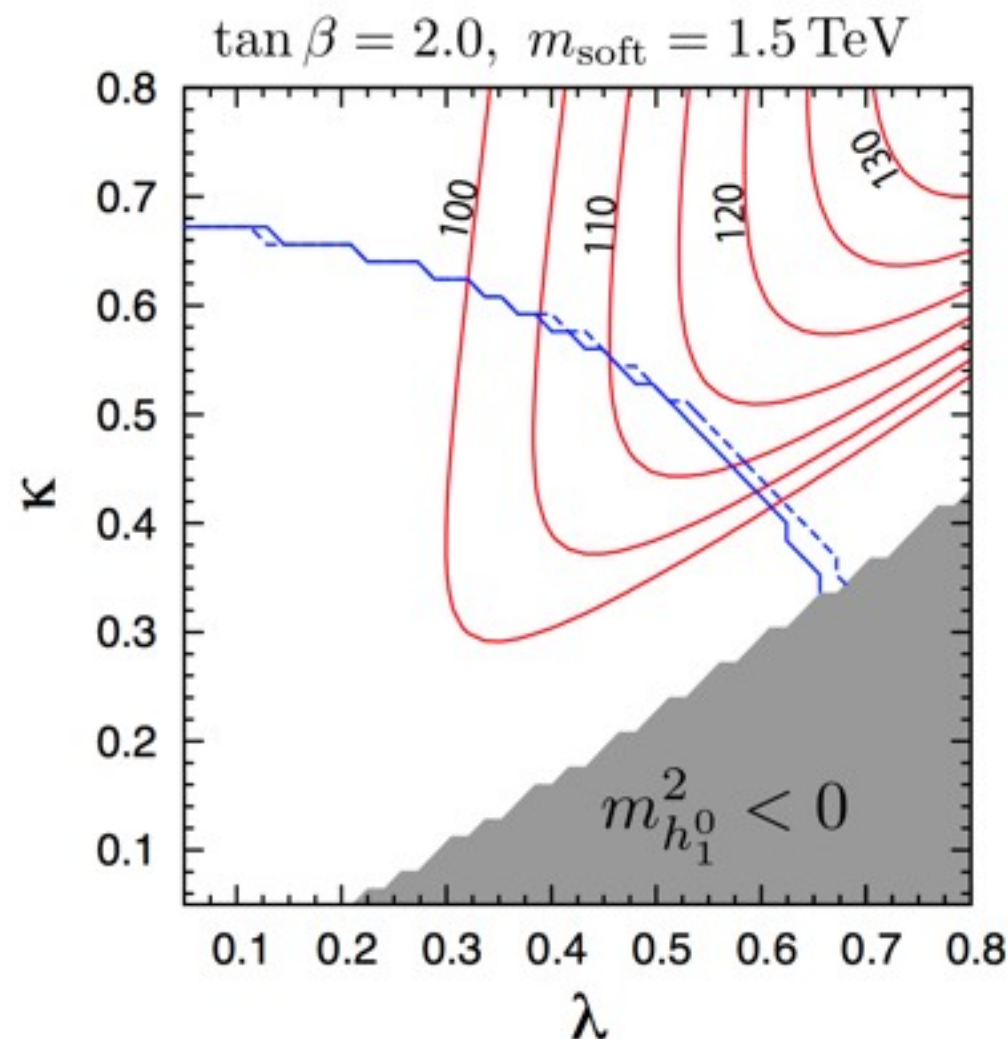
with $A_\lambda \sim 0$

[Figure from Yanagida, Yonekura, N.Y., 2012]

Most famous example is NMSSM

NMSSM \rightarrow not good in GMSB

(Difficult to realize EWSB, Landau pole, Muon $g-2$ deviation can not be explained)



$$\lambda S H_u H_d + \kappa S^3 / 3$$

$+ S q' \bar{q}'$

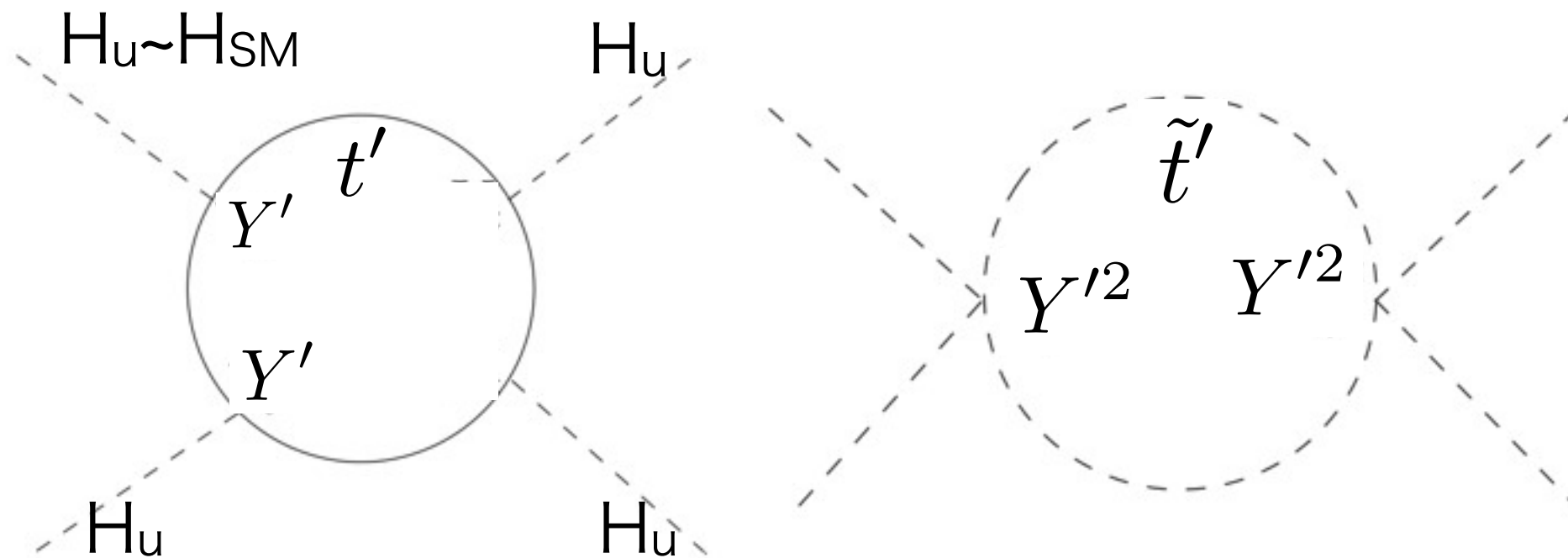
[Gouvea, Friedland, Murayama, 97']

with $A_\lambda \sim 0$

[Figure from Yanagida, Yonekura, N.Y., 2012]

Here we consider another possibility

Add top/stop like particles to raise the Higgs mass with vector-like masses



[Okada, Moroi, 92']

$10 + 10^*$ is suitable for raising the Higgs mass

$$(5+1)+(5^*+1^*)$$

$$L' \text{ Hu } N' \longleftrightarrow \text{Hu Hd S}$$

blow-up of Yukawa coupling

scalar masses are light

complete generation

$$(10+5)+(10^*+5^*) \quad Q^{*'} \text{ Hu } D'$$

blow-up of gauge coupling (since we need a messenger pair)

Model

$$W = W_{\text{MSSM}}$$

$$+ Y' Q' H_u \bar{U}' + Y'' \bar{Q}' H_d U'$$

$$+ M_{Q'} Q' \bar{Q}' + M_{U'} U' \bar{U}' + M_{E'} E' \bar{E}'$$

$10 = (Q', \bar{U}', \bar{E}')$ $\bar{10} = (\bar{Q}', U', E')$
--

In addition to the messenger sector

$$W_{\text{mess}} = (M_D + F_D \theta^2) \Psi_D \Psi_{\bar{D}} \\ + (M_L + F_L \theta^2) \Psi_L \Psi_{\bar{L}}$$

If there is no GUT breaking effect

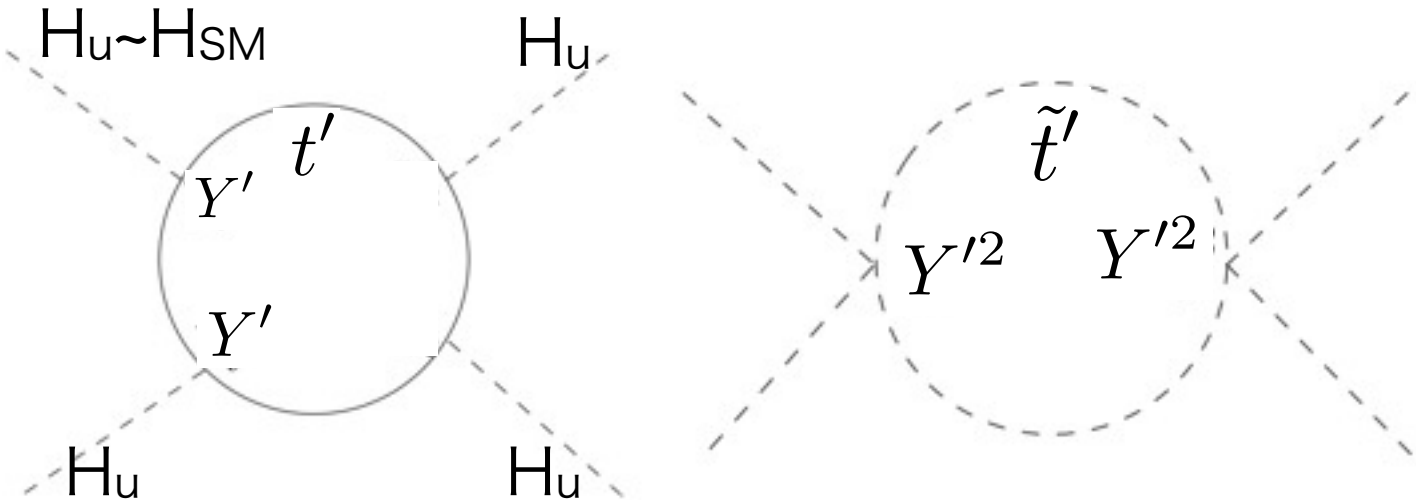
$$\frac{F_D}{M_D} \simeq \frac{F_L}{M_L} \equiv \Lambda$$

$$10 = (Q', \bar{U}', \bar{E}')$$

$$\bar{10} = (\bar{Q}', U', E')$$

$$\begin{aligned}
 W = & W_{\text{MSSM}} \\
 & + \boxed{Y' Q' H_u \bar{U}'} + Y'' \bar{Q}' H_d U' \\
 & + M_{Q'} Q' \bar{Q}' + M_{U'} U' \bar{U}' + M_{E'} E' \bar{E}'
 \end{aligned}$$

$$\Delta m_h^2 \sim \frac{3Y'^4 v^4}{4\pi^2} \ln \frac{m_S^2}{m_F^2}$$



$$10 = (Q', \bar{U}', \bar{E}')$$

$$\bar{10} = (\bar{Q}', U', E')$$

$$W = W_{\text{MSSM}}$$

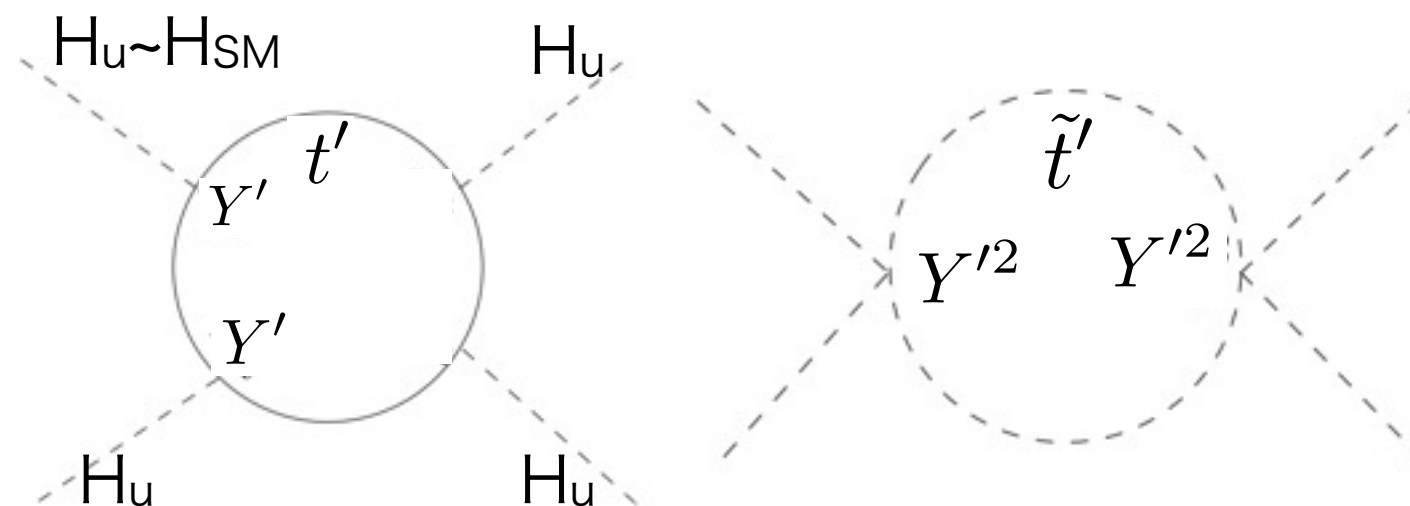
$$+ Y' Q' H_u \bar{U}' + Y'' \bar{Q}' H_d U'$$

$$+ M_{Q'} Q' \bar{Q}' + M_{U'} U' \bar{U}' + M_{E'} E' \bar{E}'$$

$$\Delta m_h^2 \sim \frac{3Y'^4 v^4}{4\pi^2} \ln \frac{m_S^2}{m_F^2}$$

m_S : scalar mass
 m_F : fermion mass

larger m_S/m_F leads to larger Higgs mass



$$10 = (Q', \bar{U}', \bar{E}')$$

$$\bar{10} = (\bar{Q}', U', E')$$

$$W = W_{\text{MSSM}}$$

$$+ Y' Q' H_u \bar{U}' + Y'' \bar{Q}' H_d U'$$

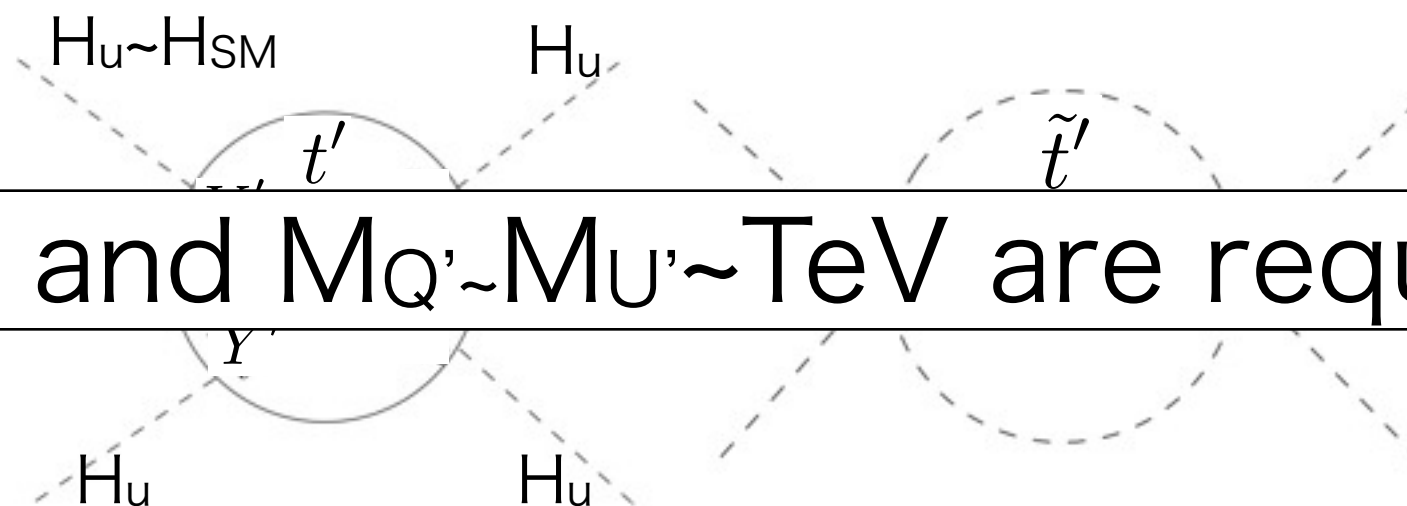
$$+ M_{Q'} Q' \bar{Q}' + M_{U'} U' \bar{U}' + M_{E'} E' \bar{E}'$$

$$\Delta m_h^2 \sim \frac{3Y'^4 v^4}{4\pi^2} \ln \frac{m_S^2}{m_F^2}$$

m_S : scalar mass
 m_F : fermion mass

larger m_S/m_F leads to larger Higgs mass

$Y' \sim 1$ and $M_{Q'} \sim M_{U'} \sim \text{TeV}$ are required



$$10 = (Q', \bar{U}', \bar{E}')$$

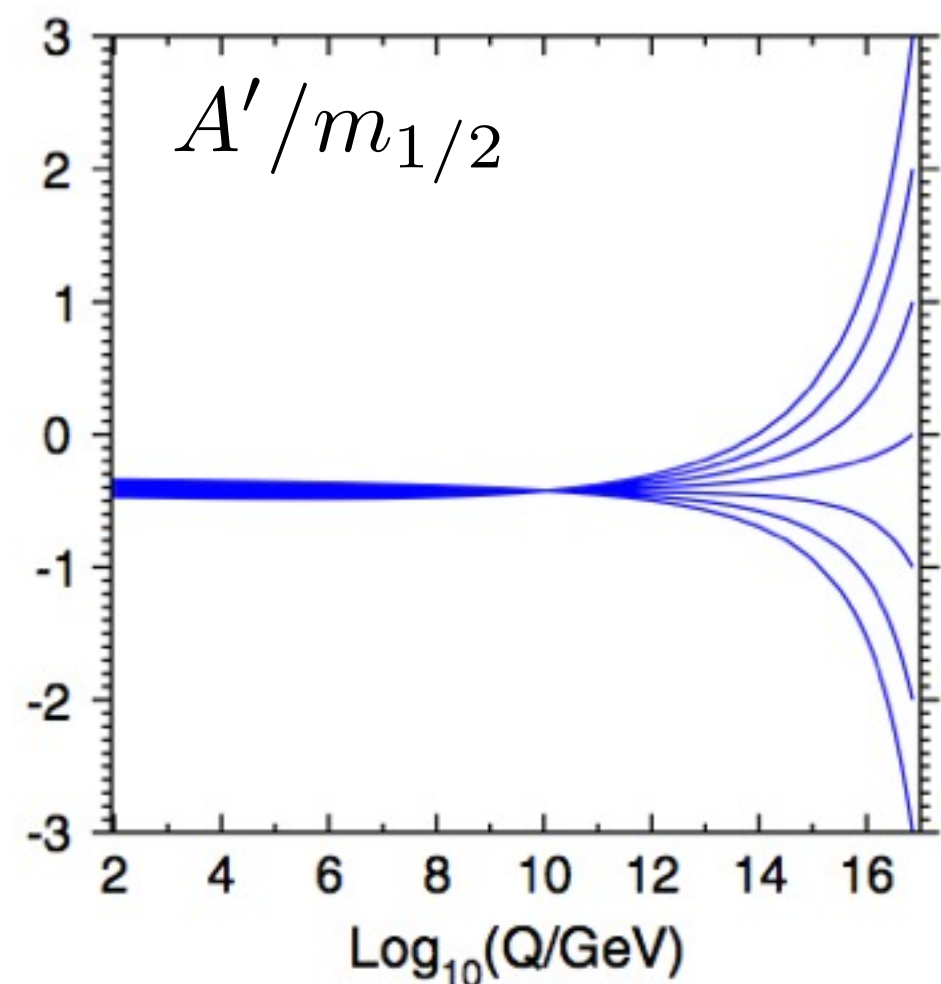
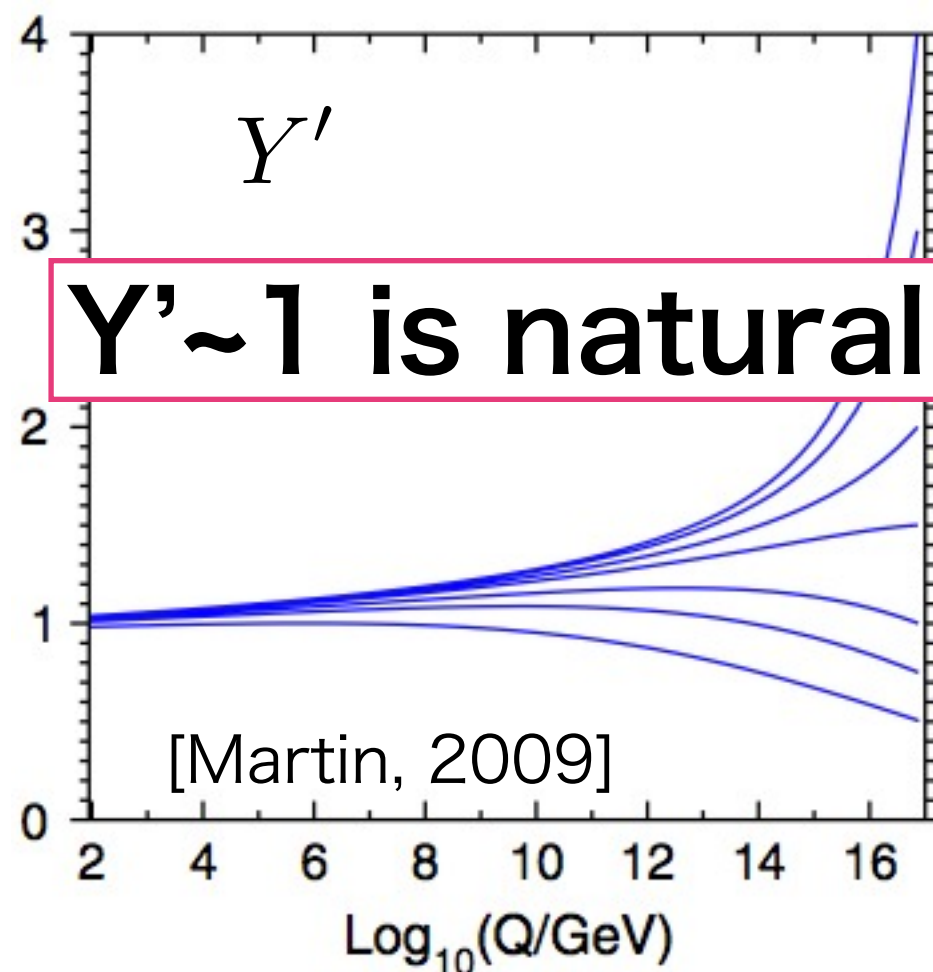
$$\bar{10} = (\bar{Q}', U', E')$$

$$W = W_{\text{MSSM}}$$

$$+ Y' Q' H_u \bar{U}' + Y'' \bar{Q}' H_d U'$$

$$+ M_{Q'} Q' \bar{Q}' + M_{U'} U' \bar{U}' + M_{E'} E' \bar{E}'$$

Y' and corresponding A parameter (and A_t) have
quasi IR fixed point

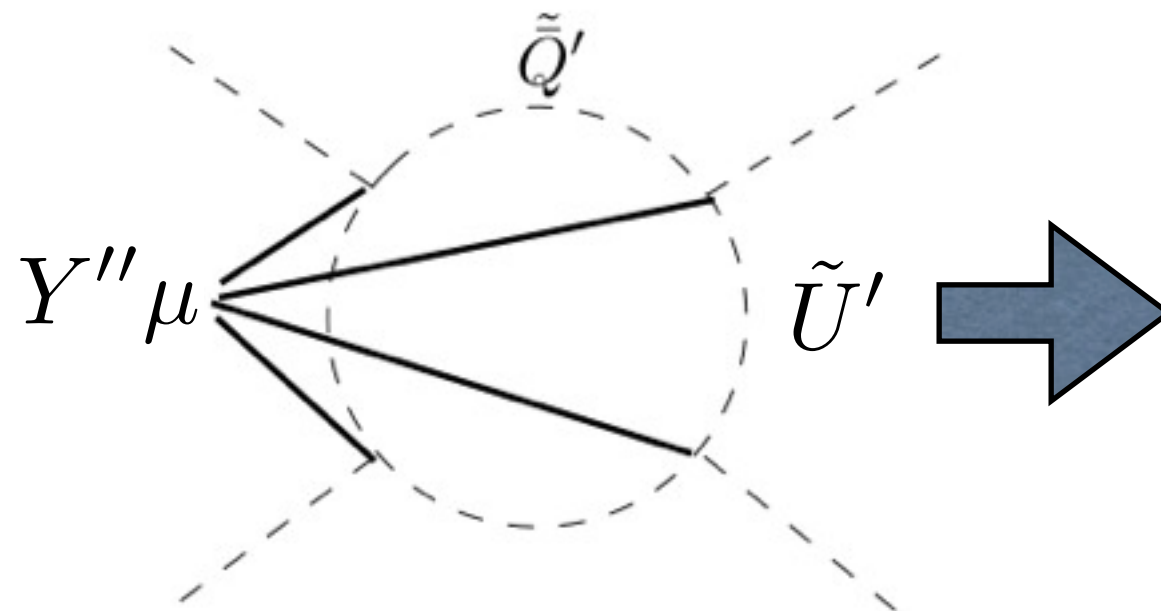


$$W = W_{\text{MSSM}}$$

$$+ Y' Q' H_u \bar{U}' + Y'' \bar{Q}' H_d U' + M_{Q'} Q' \bar{Q}' + M_{U'} U' \bar{U}' + M_{E'} E' \bar{E}'$$

$$\begin{aligned} 10 &= (Q', \bar{U}', \bar{E}') \\ \bar{10} &= (\bar{Q}', U', E') \end{aligned}$$

this coupling should be small



$$\Delta m_{h^0}^2 \simeq -\frac{3v^2}{4\pi^2} Y''^4 \sin^4 \beta \frac{\mu^4}{12M_S^4},$$

reducing the Higgs mass (similar to the sbottom loops with large LR mixing)

The correct size of μ -term, M_Q and M_U can be generated by **PQ-symmetry** breaking

$$W_{\text{ext}} = \lambda_1 \frac{\phi^2}{M_P} H_u H_d + \frac{\phi^2}{M_P} (\lambda_2 Q' \bar{Q}' + \lambda_3 \bar{U}' U' + \lambda_4 \bar{E}' E')$$

PQ breaking scale: $\langle \phi \rangle \sim 10^{10} \text{GeV} - 10^{12} \text{GeV}$

Reduced Plank mass: $M_P \simeq 2.4 \times 10^{18} \text{GeV}$

$$\begin{aligned} \mu &\sim M_Q \sim M_U \sim M_E \\ &\sim \frac{\langle \phi \rangle^2}{M_P} \sim 100 - 1000 \text{ GeV} \end{aligned}$$

The unwanted interaction is automatically suppressed by **PQ-symmetry**

Requiring that

$$W = \lambda_1 \frac{\phi^2}{M_P} H_u H_d + \frac{\phi^2}{M_P} (\lambda_2 Q' \bar{Q}' + \lambda_3 \bar{U}' U' + \lambda_4 \bar{E}' E') \\ + W_{\text{MSSM}, \mu=0} + Y' Q' H_u \bar{U}'$$

$$Y'' \bar{Q}' H_d U' \Rightarrow \Delta m_{h^0}^2 \simeq -\frac{3v^2}{4\pi^2} Y''^4 \sin^4 \beta \frac{\mu^4}{12M_S^4},$$

Small Y'' suppresses contributions to hVV couplings

The Higgs production is almost unchanged

The unwanted interaction is automatically suppressed by **PQ-symmetry**

Requiring that

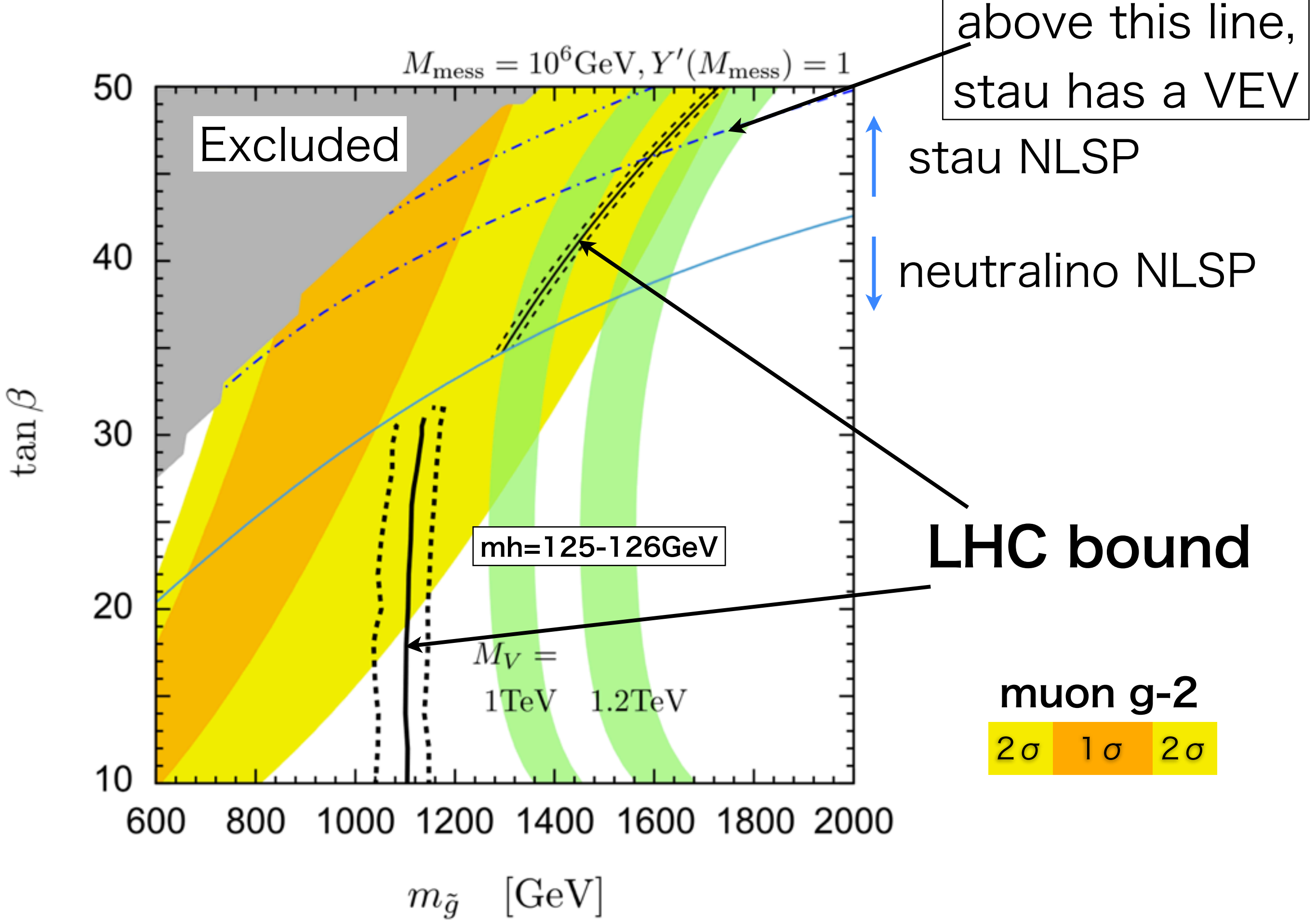
$$W = \lambda_1 \frac{\phi^2}{M_P} H_u H_d + \frac{\phi^2}{M_P} (\lambda_2 Q' \bar{Q}' + \lambda_3 \bar{U}' U' + \lambda_4 \bar{E}' E') \\ + W_{\text{MSSM}, \mu=0} + Y' Q' H_u \bar{U}'$$

~~$$Y'' \bar{Q}' H_d U' \Rightarrow \Delta m_{h^0}^2 \simeq -\frac{3v^2}{4\pi^2} Y''^4 \sin^4 \beta \frac{\mu^4}{12M_S^4},$$~~

Small Y'' suppresses contributions to hVV couplings

The Higgs production is almost unchanged

Result



Oh... it may not be attractive enough

Is it impossible to explain muon $g-2$?

Oh... it may not be attractive enough

Is it impossible to explain muon $g-2$?

Yes, it is possible

The SUSY breaking (invariant) masses of messengers can be split by GUT breaking effects (**GUT is broken!**)

$$W_{\text{mess}} = (M_D + F_D \theta^2) \Psi_D \Psi_{\bar{D}} \\ + (M_L + F_L \theta^2) \Psi_L \Psi_{\bar{L}}$$

$$\Lambda_D = \frac{F_D}{M_D} \qquad \Lambda_L = \frac{F_L}{M_L}$$

$\Lambda_D = \Lambda_L$ is not always required

We can take $\Lambda_L < \Lambda_D$

Colored and non-colored SUSY particles become more hierarchical in their masses

$$M_1 \simeq \frac{g_1^2}{16\pi^2} N_5 \left(\frac{2}{5} \Lambda_D + \frac{3}{5} \Lambda_L \right), \quad M_2 \simeq \frac{g_2^2}{16\pi^2} N_5(\Lambda_L), \quad M_3 \simeq \frac{g_3^2}{16\pi^2} N_5(\Lambda_D),$$

<div style="display: flex; flex-direction: column; align-items: center;"> <div style="margin-bottom: 10px;">↑</div> <div style="margin-bottom: 10px;">colored</div> <div style="margin-bottom: 10px;">↓</div> </div>	$m_{\tilde{Q}}^2 \simeq N_5 \frac{2}{(16\pi^2)^2} \left[\frac{4}{3} g_3^4 \Lambda_D^2 + \frac{3}{4} g_2^4 \Lambda_L^2 + \frac{3}{5} g_1^4 \left(\frac{2}{5} \Lambda_D^2 + \frac{3}{5} \Lambda_L^2 \right) \frac{1}{6^2} \right],$
	$m_{\tilde{U}}^2 \simeq N_5 \frac{2}{(16\pi^2)^2} \left[\frac{4}{3} g_3^4 \Lambda_D^2 + \frac{3}{5} g_1^4 \left(\frac{2}{5} \Lambda_D^2 + \frac{3}{5} \Lambda_L^2 \right) \left(\frac{2}{3} \right)^2 \right],$
	$m_{\tilde{D}}^2 \simeq N_5 \frac{2}{(16\pi^2)^2} \left[\frac{4}{3} g_3^4 \Lambda_D^2 + \frac{3}{5} g_1^4 \left(\frac{2}{5} \Lambda_D^2 + \frac{3}{5} \Lambda_L^2 \right) \frac{1}{3^2} \right],$
<div style="display: flex; flex-direction: column; align-items: center;"> <div style="margin-bottom: 10px;">↑</div> <div style="margin-bottom: 10px;">non-colored</div> <div style="margin-bottom: 10px;">↓</div> </div>	$m_{\tilde{L}}^2 \simeq N_5 \frac{2}{(16\pi^2)^2} \left[\frac{3}{4} g_2^4 \Lambda_L^2 + \frac{3}{5} g_1^4 \left(\frac{2}{5} \Lambda_D^2 + \frac{3}{5} \Lambda_L^2 \right) \frac{1}{2^2} \right],$
	$m_{\tilde{E}}^2 \simeq N_5 \frac{2}{(16\pi^2)^2} \left[\frac{3}{5} g_1^4 \left(\frac{2}{5} \Lambda_D^2 + \frac{3}{5} \Lambda_L^2 \right) \right],$
	$m_{H_u}^2 = m_{H_d}^2 = m_{\tilde{L}}^2.$

Colored and non-colored SUSY particles become more hierarchical in their masses

$$M_1 \simeq \frac{g_1^2}{16\pi^2} N_5 \left(\frac{2}{5} \Lambda_D + \cancel{\frac{3}{5} \Lambda_L} \right), \quad M_2 \simeq \frac{g_2^2}{16\pi^2} \cancel{N_5(\Lambda_L)}, \quad M_3 \simeq \frac{g_3^2}{16\pi^2} N_5(\Lambda_D),$$

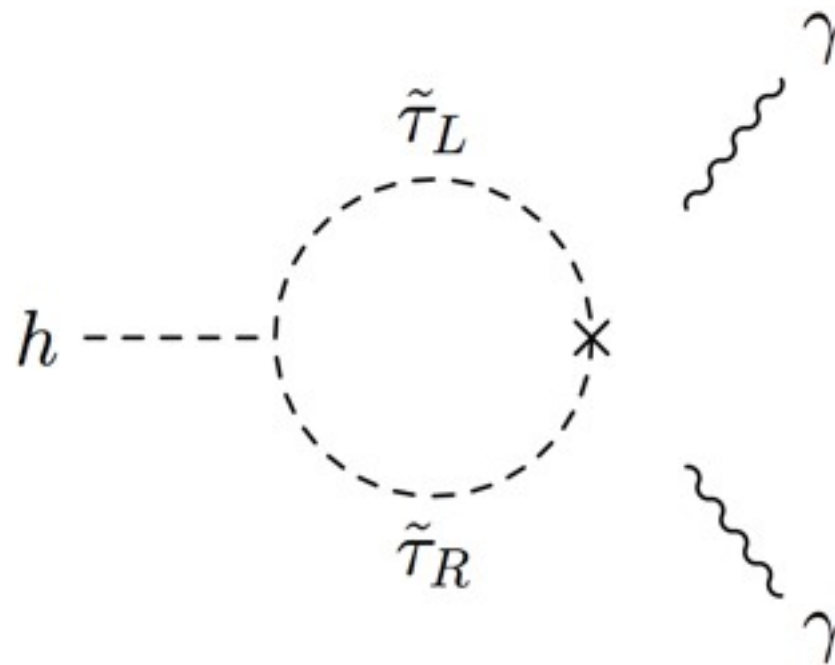
<div style="display: flex; flex-direction: column; align-items: center;"> <div style="margin-bottom: 10px;">↑</div> <div style="margin-bottom: 10px;">colored</div> <div style="margin-bottom: 10px;">↓</div> </div>	$m_{\tilde{Q}}^2 \simeq N_5 \frac{2}{(16\pi^2)^2} \left[\frac{4}{3} g_3^4 \Lambda_D^2 + \cancel{\frac{3}{4} g_2^4 \Lambda_L^2} + \frac{3}{5} g_1^4 \left(\frac{2}{5} \Lambda_D^2 + \cancel{\frac{3}{5} \Lambda_L^2} \right) \frac{1}{6^2} \right],$
	$m_{\tilde{U}}^2 \simeq N_5 \frac{2}{(16\pi^2)^2} \left[\frac{4}{3} g_3^4 \Lambda_D^2 + \frac{3}{5} g_1^4 \left(\frac{2}{5} \Lambda_D^2 + \cancel{\frac{3}{5} \Lambda_L^2} \right) \left(\frac{2}{3} \right)^2 \right],$
	$m_{\tilde{D}}^2 \simeq N_5 \frac{2}{(16\pi^2)^2} \left[\frac{4}{3} g_3^4 \Lambda_D^2 + \frac{3}{5} g_1^4 \left(\frac{2}{5} \Lambda_D^2 + \cancel{\frac{3}{5} \Lambda_L^2} \right) \frac{1}{3^2} \right],$
<div style="display: flex; flex-direction: column; align-items: center;"> <div style="margin-bottom: 10px;">↑</div> <div style="margin-bottom: 10px;">non-colored</div> <div style="margin-bottom: 10px;">↓</div> </div>	$m_{\tilde{L}}^2 \simeq N_5 \frac{2}{(16\pi^2)^2} \left[\cancel{\frac{3}{4} g_2^4 \Lambda_L^2} + \frac{3}{5} g_1^4 \left(\frac{2}{5} \Lambda_D^2 + \cancel{\frac{3}{5} \Lambda_L^2} \right) \frac{1}{2^2} \right],$
	$m_{\tilde{E}}^2 \simeq N_5 \frac{2}{(16\pi^2)^2} \left[\frac{3}{5} g_1^4 \left(\frac{2}{5} \Lambda_D^2 + \cancel{\frac{3}{5} \Lambda_L^2} \right) \right],$
	$m_{H_u}^2 = m_{H_d}^2 = m_{\tilde{L}}^2.$

Colored and non-colored SUSY particles become
more hierarchical in their masses

Muon $g-2$ is easier to be explained

Colored and non-colored SUSY particles become more hierarchical in their masses

Muon $g-2$ is easier to be explained
Stau-loop enhances $H \gamma \gamma$ signal with large LR mixing



[Carena, Gori, Shah, Wagner, 2012]

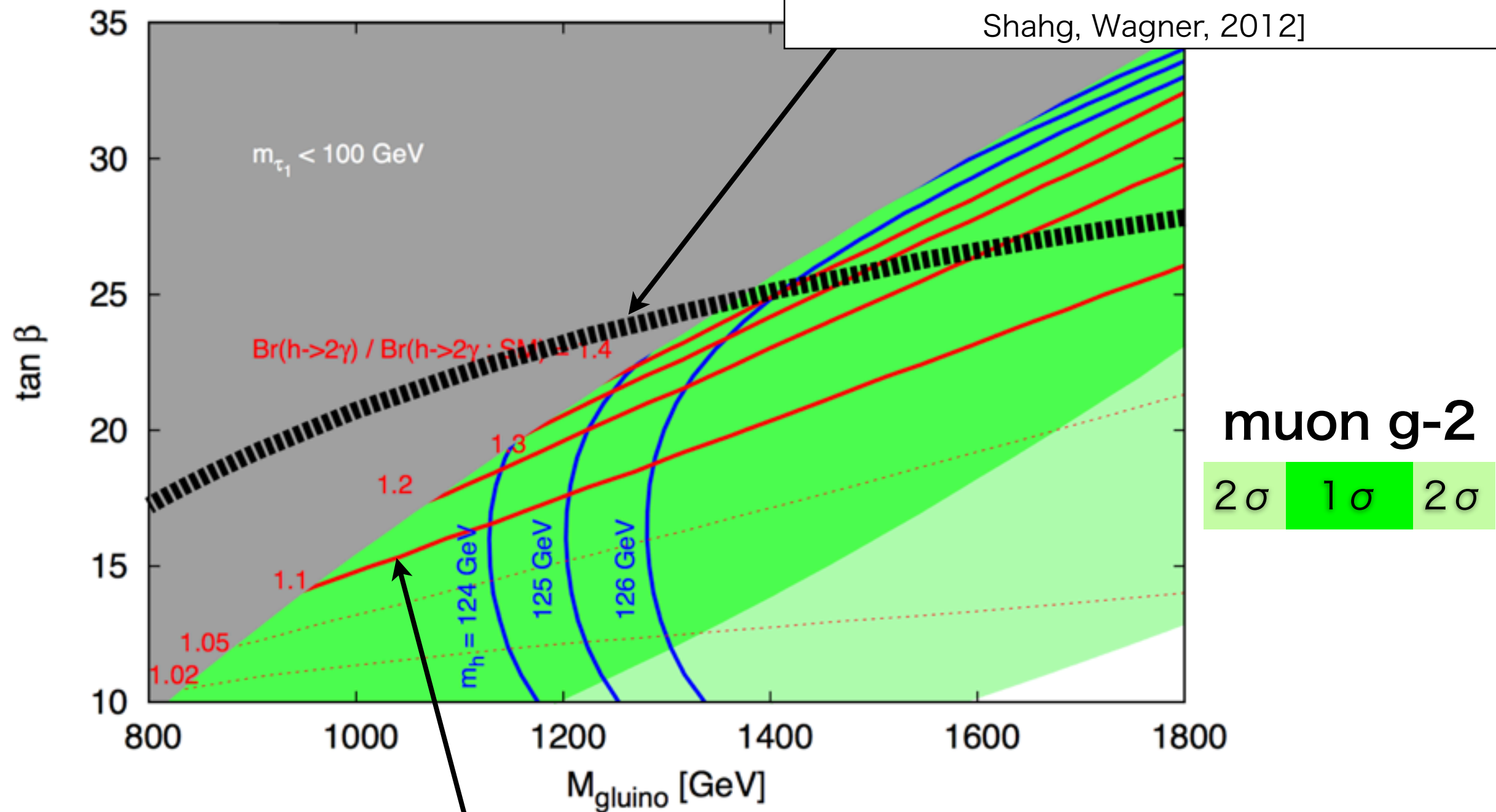
Result 2

$$M_Q = M_U = 900 \text{ GeV}, \quad \Lambda_L / \Lambda_{\bar{D}} = 0.4$$

$$M_{\text{mess}} = 4 \times 10^5 \text{ GeV}$$

Vacuum stability bound

[Hisano, Sugiyama, 2010v2; Carena, Gori, Low, Shahg, Wagner, 2012]



$$\text{Br}(h \rightarrow \gamma\gamma) / \text{Br}(h \rightarrow \gamma\gamma)_{\text{SM}}$$

[Sato, Tobioka, N.Y, 2012]

Summary

- SUSY SM + vector-like matters in gauge mediation is attractive
- Higgs boson mass, muon $g-2$ and enhanced di-photon rates (optional) are all explained
- Consistent with perturbative GUT
- Not very large fine-tuning is required compared to high scale SUSY breaking scenarios

Thank you